ENERGY DECAYS LOCALLY EVEN IF TOTAL ENERGY GROWS ALGEBRAICALLY WITH TIME

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0. Introduction. In this note we announce energy decays locally like $t^{-2+\kappa}$ for solutions of hyperbolic equations, with coefficients that depend upon both position and time, in the exterior of star-shaped domains in \mathbb{R}^3 . Here κ is a positive constant, depending on the coefficients, defined explicitly by (8) below. Our results generalize those of Zachmanoglou [4]. He considered a class of equations with time-independent coefficients (see (12) below) and proved under hypotheses roughly analogous to ours that in \mathbb{R}^n ($n \ge 3$) energy decays locally like $t^{-1+\mu}$ ($1 > \mu \ge 0$). A more important difference between the equations we consider here and those considered by Zachmanoglou in [4] is that we treat equations with solutions whose total energy may grow algebraically with t while the total energy of solutions of the equations with bounded total energy decays locally like t^{-2} , but under more stringent hypotheses than those used here.

We now set the scattering problem whose solutions we investigate. Let V be the exterior of a closed, bounded subset B of \mathbb{R}^3 , and let n be the outward unit normal to ∂B . We assume that the origin lies interior to B and that $\partial V \equiv \partial B$ is star-shaped:

(1)
$$\min_{x\in\partial V}\frac{n\cdot x}{r}\geq 0,$$

where $x = (x_1, x_2, x_3)$ and $r^2 = x \cdot x$. Let $Q = (V \cup \partial V) \times [0, \infty)$. We use the notation $\nabla^{(3)} = (\partial/\partial x_1, \partial/\partial x_2, \partial/\partial x_3), \nabla = (\nabla^{(3)}, \partial/\partial t)$. We take as given a symmetric 3×3 matrix $E, 1 \times 3$ matrices a and b, and functions c and d which satisfy the following hypothesis:

(Hypothesis H₁) (a) b, c, and E are in $C^1(Q)$; a and d are in $C^2(Q)$, (b) for some $d_0 > 0$, $d(x, t) \ge d_0$ if $(x, t) \in Q$.

Let the transpose of a matrix M (or m) be M^T (or m^T). We suppose that E is uniformly elliptic in Q, namely that there exist positive constants c_0 and

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 C_0 such that

$$C_0 \geq \max_{|\xi|=1} \xi E \xi^T \geq \min_{|\xi|=1} \xi E \xi^T \geq c_0$$

Finally we adopt the notation

$$A = \begin{pmatrix} E & a^T \\ a & -d \end{pmatrix}, \qquad D = \begin{pmatrix} E_t & b^T \\ b & -c \end{pmatrix},$$

and

$$(\hat{\cdot}) = \frac{(\cdot)}{\min_{|\xi|=1} \xi E \xi^T}.$$

We consider solutions of the mixed initial boundary value problem

(2)
$$\nabla (A(\nabla u)^T) + (b - a_t) \cdot \nabla^{(3)} u + \frac{1}{2}(d_t - c)u_t = 0$$
 $(x \in V, t > 0),$

(3) $u(x, t) = 0 \qquad (x \in \partial V, t \ge 0),$

(4)
$$u(x, 0) = f(x), u_t(x, 0) = g(x) \quad (x \in V),$$

where f and g are functions in $C^1(V \cup \partial V)$ with compact support.

1. Norms and constants. We use the following norms:

$$N(\cdot) \equiv \max_{V \cup \partial V} |\cdot|$$
 and $N'(\cdot) \equiv \max_{r \ge 0} N(\cdot)$.

We shall assume that the coefficients c, d and the matrix coefficients a, b, and E are such that the following are positive numbers:

(5)

$$\begin{aligned}
\alpha_{1} &= 2N'(rd_{r}/d) + 2N'(ra_{r}/d) + N'(r(d_{t} - c)/d), \\
\alpha_{2} &= 2N'(r\hat{E}_{r}) + 2N'(r\hat{a}_{r}) + 4N'(r(\hat{a}_{t} - \hat{b})) + N'(r(\hat{d}_{t} - \hat{c})) \\
&+ 8N(r^{-2})N'[r^{4}(\frac{1}{2}(d_{t} - c)_{t} - \nabla^{(3)} \cdot (a_{t} - b))]N'(\hat{1}), \\
\alpha_{1}' &= N'(tb/d) + N'(tc/d), \quad \alpha_{2}' = N'(t\hat{E}_{t}) + N'(t\hat{b}).
\end{aligned}$$

In addition to the smoothness conditions on the coefficients already imposed, this hypothesis imposes decay rates on the coefficients in (1) and on their time and space derivatives.

Let Ω be defined by the equation

(6)
$$\begin{cases} 1 + N'(\hat{d}) + N'(r^{-2})N'(\hat{1})[4N'(r^{3}d_{r}) + 2N'(r^{3}(d_{t} - c))] \\ + 4(N'(r^{-4}))^{1/2}N'(\hat{1})N'(r^{2}a) + 2N'(\hat{a})\}\Omega = \frac{1}{2}. \end{cases}$$

Again we shall assume that the coefficients of (1) are such that the above norms are well defined.

For any $\varepsilon \in (0, 1)$, each T > 0, and some small $\delta > 0$, we define

(7)
$$\mathscr{D}(T) = \{x \mid r \leq \varepsilon \Omega T\} \cap V,$$

(8)
$$\kappa = (1 - \varepsilon)^{-1} [\max(\alpha_1, \alpha_2) + (1 + (\varepsilon \Omega)^2) \max(\alpha'_1, \alpha'_2)]$$

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(9)
$$s = \max[N'(t\hat{E}_t) + N'(t\hat{b}), N'(tb/d) + N'(tc/d)],$$

and

$$(10) q = -1 + \delta + s.$$

Note that if the differential equation (2) is the wave equation, then $\kappa = 0$.

2. Local energy decay. We make the following major hypothesis:

(Hypotheses H_2) The N'-norm of each of the following is a positive number:

$$r^{3+q}(E-dI), \quad r^{4+q}E_r, \quad r^{4+q}d_r, \quad r^{4+q}(d_t-c), \quad r^{3+q}a,$$

$$r^{4+q}a_r, \quad r^{4+q}(a_t-b), \quad r^{4+q}[2(d_t-c)+d_t-2\nabla^{(3)}\cdot a],$$

$$r^{5+q}(d_t-c)_t - 2\nabla^{(3)}\cdot (a_t-b), \quad tE_t, r^{4+q}E_t, tb, \text{ and } t^{4+q}c.$$

Let $\mathscr{E}_{loc}(x, t) = \frac{1}{2} \int_{\mathscr{D}} \left[du_t^2 + \nabla^{(3)} u E(\nabla^{(3)} u)^T \right] dx$, and let $\mathscr{E}(x, 0)$ be the total initial energy associated with u.

THEOREM 1. Suppose that ∂V is star-shaped, E is uniformly strongly elliptic, Hypothesis H_1 is satisfied and the initial data (f, g) in (4) are smooth and have compact support in V. Then the unique solution to Problems (2)-(4) has compact support. Moreover, suppose $T_0 \equiv T_0(\varepsilon)$ is so large that $\partial V \subset \mathcal{D}(T_0)$, where \mathcal{D} is defined by (7) and Hypothesis H_2 is satisfied by the coefficients of the differential equation (2). Then, for each $\varepsilon \in (0, 1)$, there exist positive constants M, K, and κ , with κ defined by (8), such that for $T > T_0$

(11)
$$\mathscr{E}_{loc}(x,T) \leq \frac{K\mathscr{E}(x,0)}{T^2} \left\{ 1 + \left(\frac{T}{T_0}\right)^{\kappa} e^{M/T_0} \left[1 - \left(\frac{T_0}{T}\right)^{\kappa} + \frac{M}{(1+\kappa)T_0} \right] \right\}.$$

The theorem holds even if for each positive number p there exists a t > p for which the quadratic form associated with the matrix D defined by (1) fails to be negative semidefinite on $V \cup \partial V$. The negative semidefiniteness of D was a major hypothesis in [1].

COROLLARY 1. If the quadratic form associated with D is negative semidefinite on $(V \cup \partial V) \times [T_0, \infty)$, then the energy decay estimate (11) holds with α'_1 and α'_2 , which enter into the definition of κ , both zero and with the exponent q in Hypothesis H_2 equal to $-1 + \delta$.

In the case of the wave equation our decay estimate reduces to $\mathcal{O}(t^{-2})$, which is the same as that obtained by C. S. Morawetz in [2]. E. C. Zachmanoglou [4] proved energy decays locally like $t^{-1+\mu}$ ($1 > \mu \ge 0$) for solutions of hyperbolic equations of the form

(12)
$$\nabla^{(3)}(E(x)(\nabla^{(3)}u)^T) - c(x)u - d(x)u_{tt} = 0$$

by generalizing the argument used by Morawetz in [3] in treating the wave equation. To establish a faster rate of energy decay as $t \to \infty$ than the rate Zachmanoglou establishes in [4], we have to impose Hypothesis H₂, which is more stringent by a factor of r than his analogous conditions. But the total energy of solutions of equations of the form (2) satisfying Hypothesis H₂ may grow algebraically with t. In [1] we considered equation (2) with E = E(x, t), c = 0, and d = 1 under the hypotheses that $E_t \leq 0$, $c_0 \geq 1$, and for $t \geq N$ and $r \geq \varepsilon t + c$, $|r\nabla^{(3)}E| = \mathcal{O}(t^{-2-\delta})$, $|x(E - I)| = \mathcal{O}(t^{-1-\delta})$, and $|rE_r| = \mathcal{O}(t^{-2-\delta})$ for some positive c and δ .

Our methods of proof in [1] and of Theorem 1 are similar; the estimates we use to prove Theorem 1 are much sharper. Both results are based on the divergence identity

$$\nabla [(\alpha \cdot \nabla u)(Aw) - \alpha^{T}(\nabla uAw)/2 + u(\gamma A + B)w + C^{T}u^{2}/2]$$

$$\equiv [\alpha \cdot \nabla u + \gamma u](\nabla (Aw)) + (\nabla \cdot C)u^{2}/2 + u[C + (\nabla \gamma)A + \nabla \cdot B] \cdot \nabla u$$

$$+ \nabla u[\gamma A - (\nabla \cdot \alpha)A/2 - (\alpha \cdot \nabla)A/2 + \nabla \alpha A]w,$$

where *B* is an antisymmetric 4×4 matrix with $B^{i^*} = 0$ (i = 1, 2, 3), $w = (\nabla u)^T$, $B^{4^*} = [-(\beta + t^2 r^{-2})x, 0]d$, $\alpha = (2xt, r^2 + t^2)$, $\gamma = 2t$, $C = (2tx dr^{-2} + \Delta d, -d - t^2 dr^{-2})$, β is a solution of

$$(d\beta)_r + 3(d\beta)r^{-1} = 3 dr^{-1} + t(d_t - c)r^{-1} + t^2 d_r r^{-2},$$

 $\nabla \alpha = (\alpha_{x_i}^j)$, and $\Delta d = 2t(b - a_i) - 2a + [(\beta d)_t + (t^2 d)_t r^{-2}]x - 2tx dr^{-2}$. We integrate this identity over the space-time domain bounded by the planes $t = T_0$, $t = T(T > T_0)$, and the surface $\partial V \times [T_0, T]$, apply the divergence theorem, and estimate carefully to prove Theorem 1.

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